

WRF Version 2: Physics Update



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WRF Physics



- Diffusion
- Radiation (longwave and shortwave)
- Surface (surface layer and land-surface)
- Planetary Boundary Layer
- Cumulus Parameterization
- Microphysics

New Options in Version 2

- Grell-Devenyi Ensemble Cumulus Scheme
- WRF Single-Moment Microphysics (3, 5, 6-class options)
- Noah Land Surface Model
- RUC Land Surface Model
- Yonsei University Planetary Boundary Layer Scheme

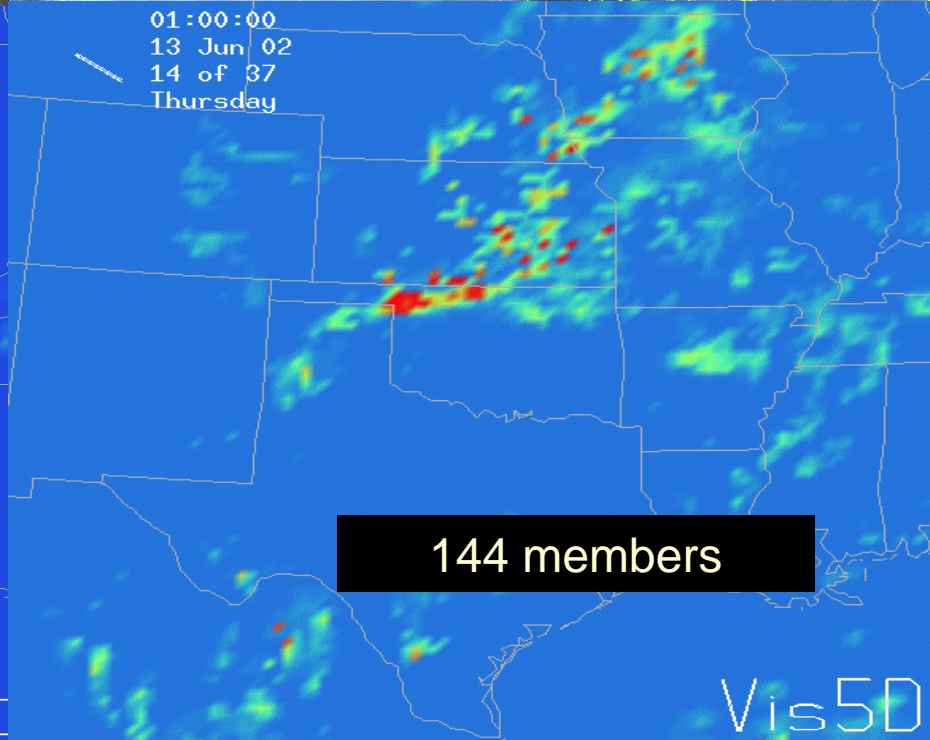
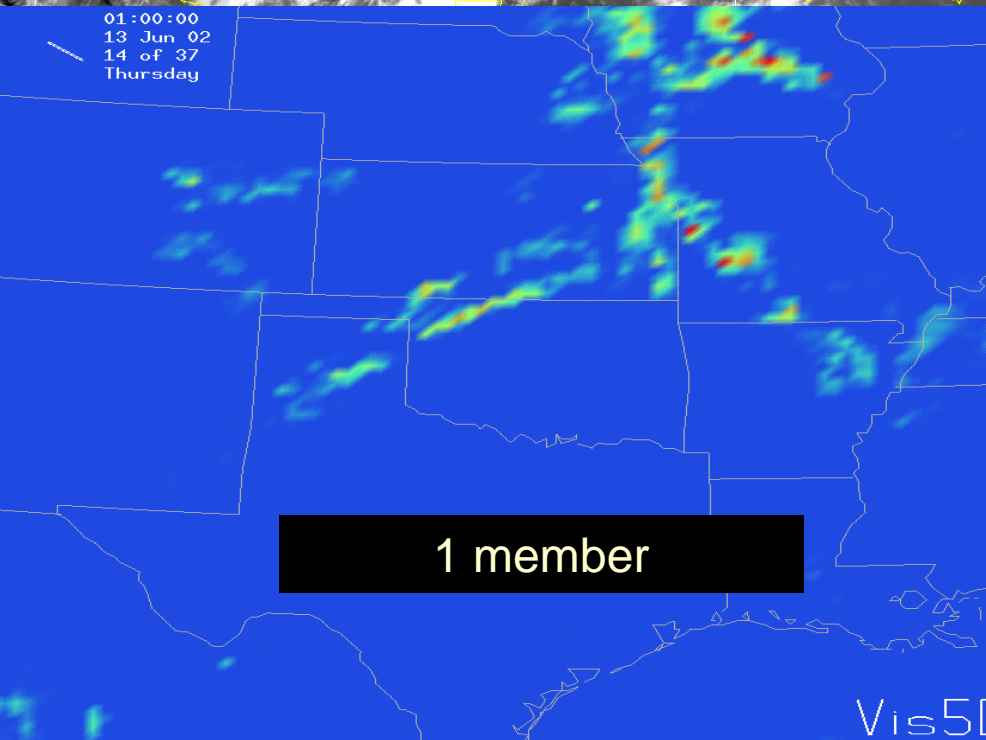
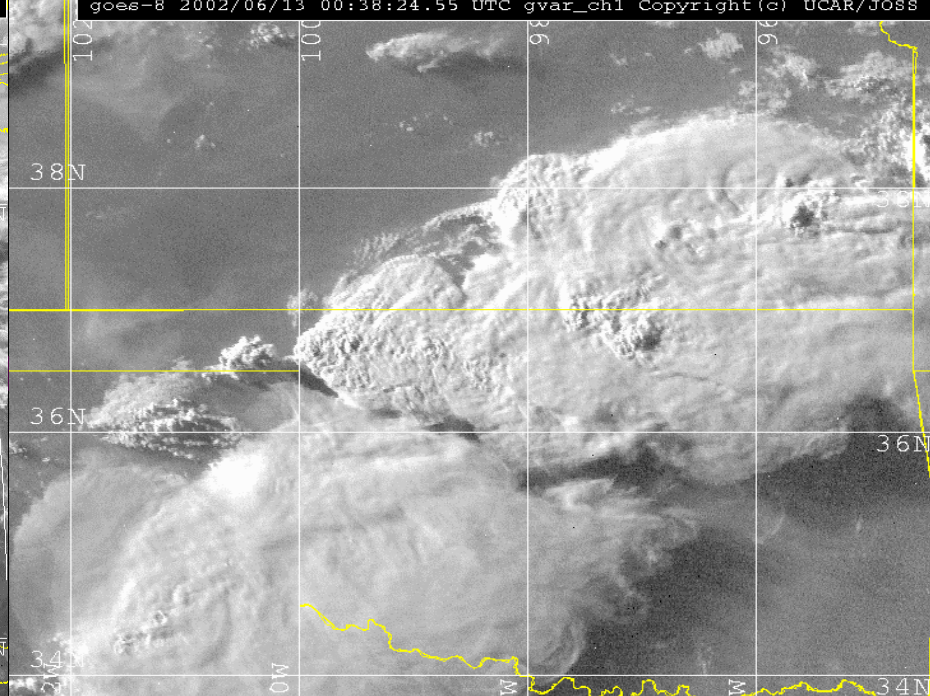
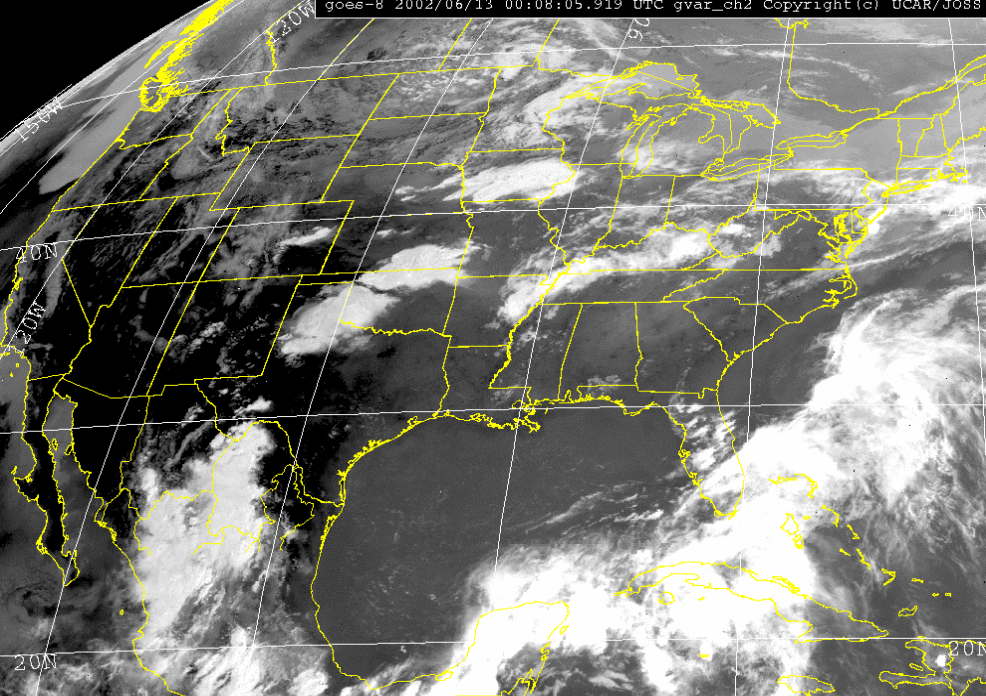
Grell-Devenyi Ensemble Cumulus Parameterization

Developers: Georg Grell, Dezso Devenyi
(NOAA/FSL)

- Typically 144 ensemble members calculated per grid column (still efficient)
- Members differ in
 - Closure (CAPE, dCAPE/dt, moisture conv, etc)
 - Trigger (maximum cap strength)
 - Precipitation efficiency
 - Other parameters (updraft assumptions, etc.) may be varied

Grell-Devenyi Ensemble Cumulus Parameterization

- Currently feedback (precip, heating, moistening profiles) is just ensemble average with equal weights
- Statistical methods can be used to train weights regionally and/or diurnally
- Scheme is potentially more optimizable than individual schemes



WRF Single-Moment Microphysics

Developers: Song-You Hong, Jimmy Dudhia,
Shuhua Chen, Jeong-Ock Lim

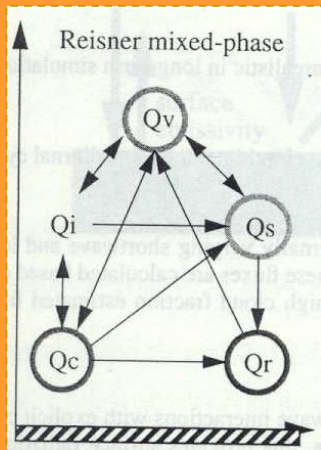
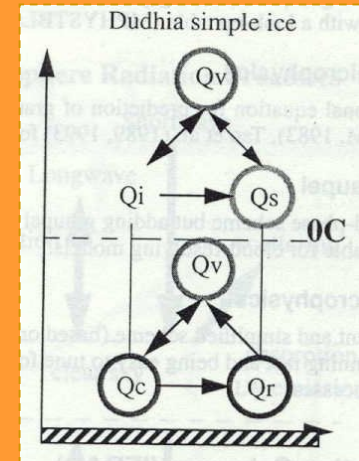
3 schemes

- 3-class (cloud/ice, snow/rain, vapor)
- 5-class (cloud, ice, snow, rain, vapor)
- 6-class (cloud, ice, snow, rain, graupel, vapor)

Bulk parameterization

- Simple ice (Dudhia, 1989), WSM3
: 3 arrays of moisture

➔ q_v, q_{ci}, q_{rs}



- Mixed phase (Reisner et al., 1998), WSM5
: 5 arrays of moisture

➔ q_v, q_c, q_i, q_r, q_s

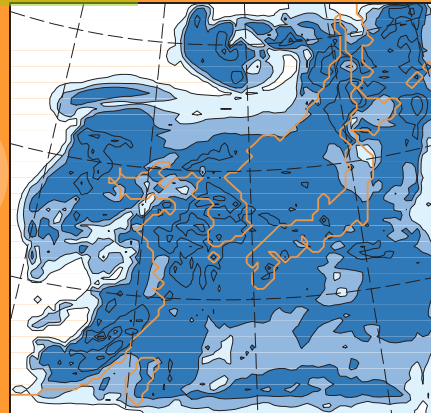
WRF Single-Moment Microphysics

- WSM3 and WSM5 based on Dudhia (1989) and MM5 Reisner '1' schemes
- WSM6 adds graupel, modified from Lin et al.
- Schemes are distinguished from older schemes mostly by ice crystal size/number assumptions

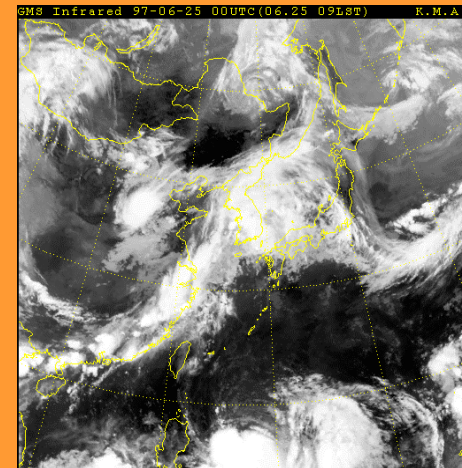
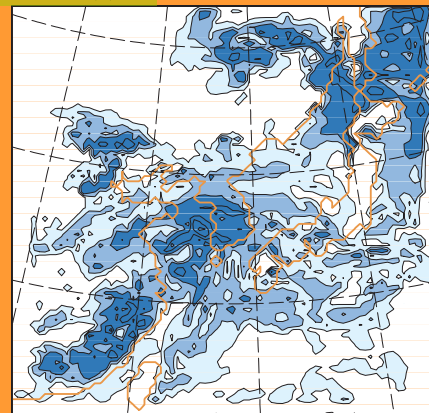
Major modifications suggested by Hong et al. (2004)

	(Rutledge and Hobbs, 1983)	(Hong et al, 2004)
Number concentration of cloud ice	$N_I (m^{-3}) = 10^{-2} \exp[0.6(T_0 - T)]$	$N_I = c(\rho q_I)^d$
Ice nuclei number	$N_I (m^{-3}) = 10^{-2} \exp[0.6(T_0 - T)]$	$N_{I0} = 10^3 \exp[0.1(T_0 - T)]$
Intercept parameter for snow	$N_{0S} = 2 \times 10^7 m^{-4}$	$N_{0S} (m^{-4}) = 2 \times 10^6 \exp\{0.12(T_0 - T)\}$

old



new



23 -25 June 1997


Heavy Rainfall Case

Ice crystal property

(Mass, Diameter, Mixing ratio, Ice number)

$$V_I (ms^{-1}) = 3.29(\rho q_I)^{0.16} \quad : \text{Heymsfield and Donner(1990)} \quad (\text{HD1990})$$

$$V_I = xD^y, m = \alpha D^\beta \quad : \text{Heymsfield and laquinta (2000)} \quad (\text{HI2000})$$

$$mN_i = \rho q_i$$


$$N_i = c(\rho q_i)^d$$


$$V_I (ms^{-1}) = 1.49 \times 10^4 D^{1.31},$$

$$D(m) = 11.9m^{0.5}$$

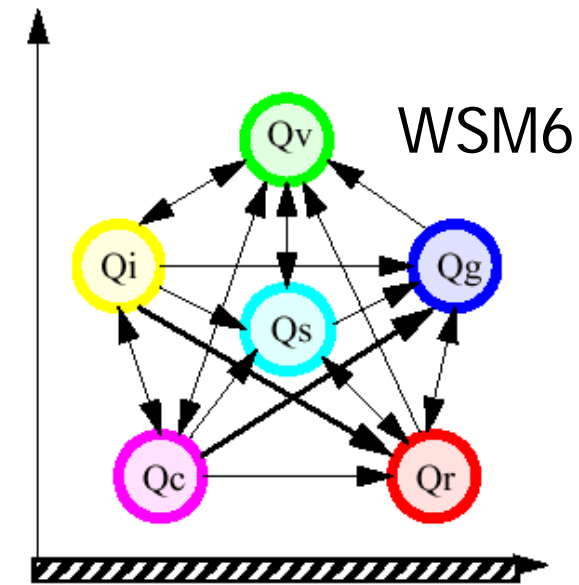
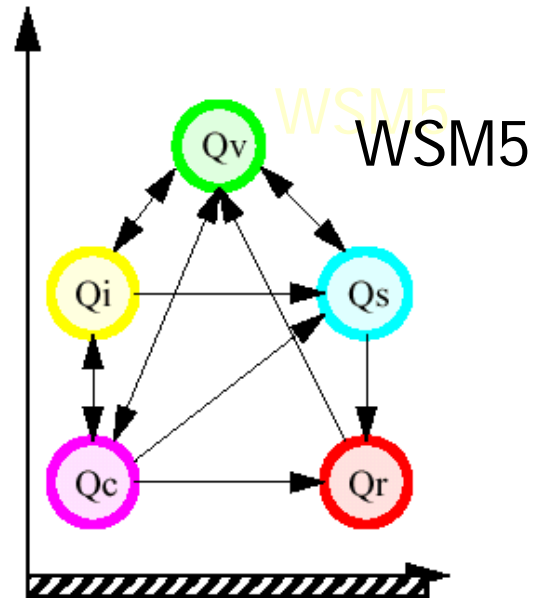
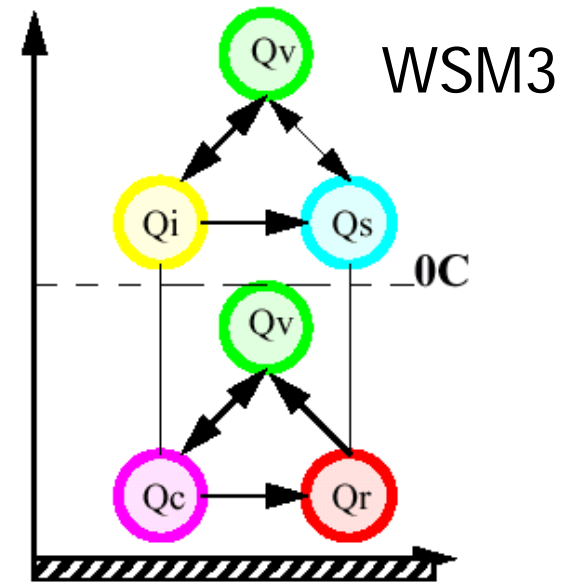
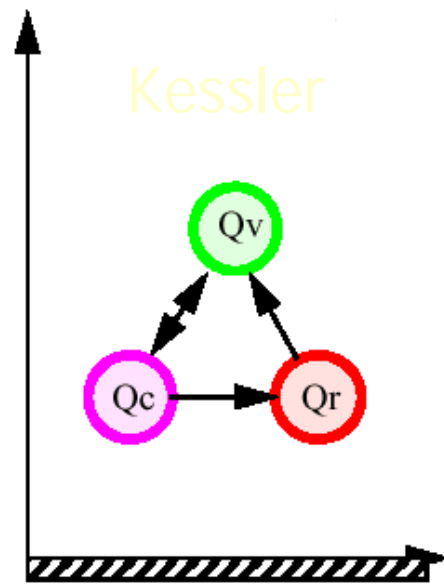
$$N_I (m^{-3}) = 5.38 \times 10^7 (\rho q_i)^{0.75}$$

$$\rho q_I (kgm^{-3}) = 4.92 \times 10^{-11} N_I^{1.33}$$

Development of WSM6

- Riming and graupel processes
- Accounts for relative fall speeds in accretion (idea from M. Gilmore)
- Incorporates melting into fall sub-steps
- Calculation order: sensitivity to timestep length minimized
- Comparison with Lin, Farley and Orville (LFO)

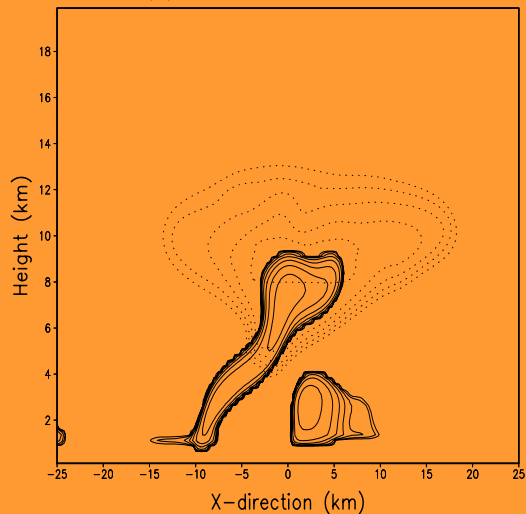
Illustration of Microphysics Processes



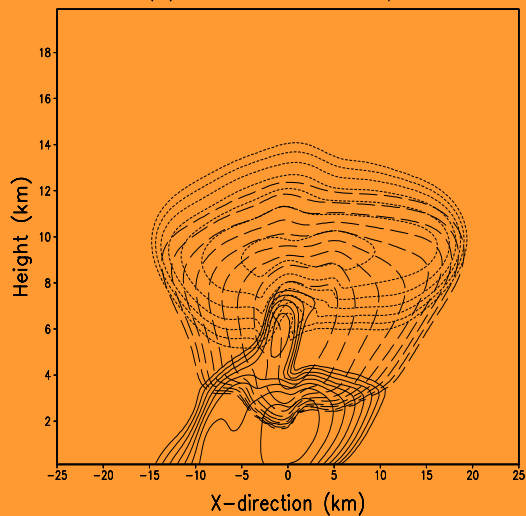
Development of WSM6

WSM6

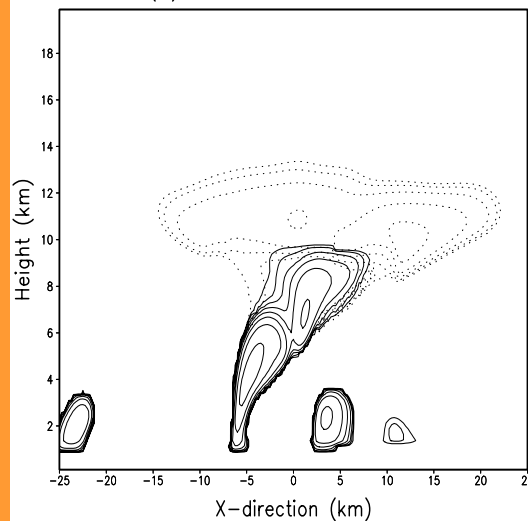
(c) Cloud and ice waters



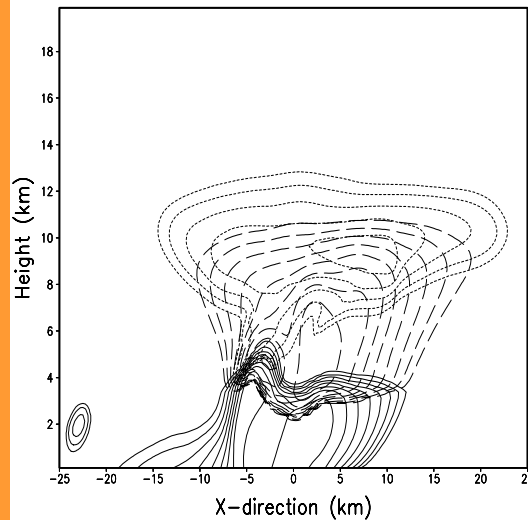
(d) Rain, Snow, Graupel



(c) Cloud and ice waters



(d) Rain, Snow, Graupel



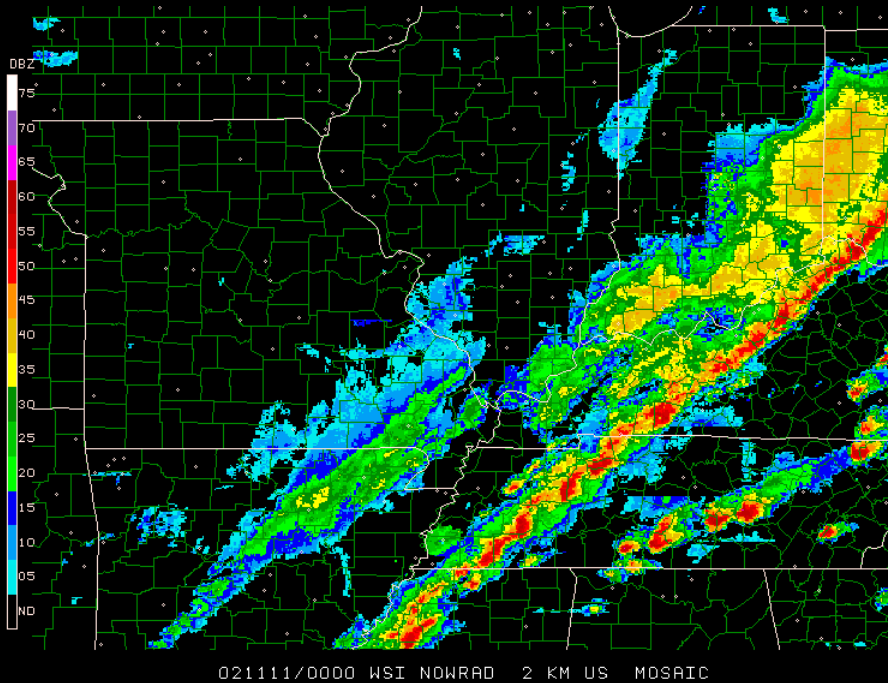
LFO

Real-Data case

- 10-11 November 2002 tornado outbreak
- East-central US
- 4 km cloud-resolving simulation
- 12 hour forecast
- Simulated reflectivity from
 - WSM6
 - Lin et al (LFO)

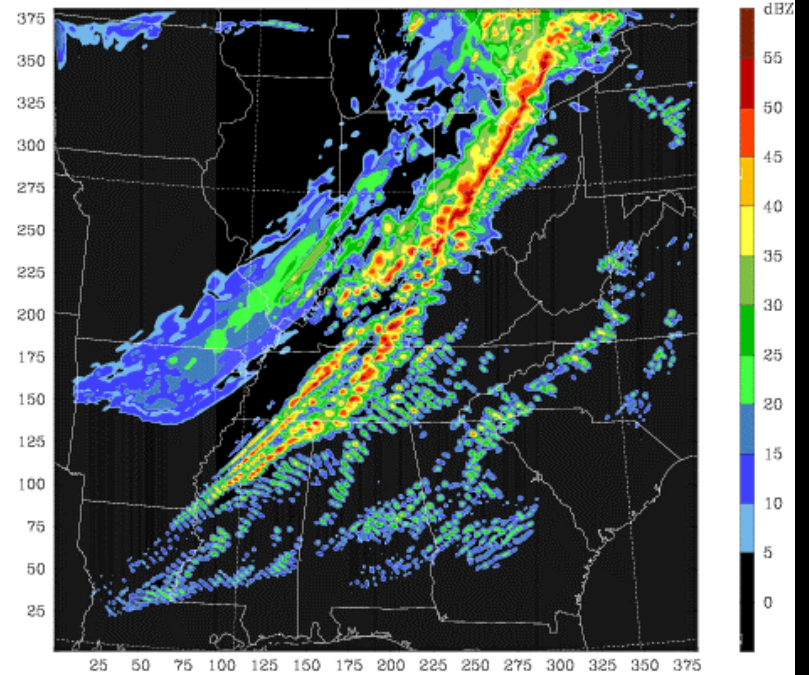
00Z 11 Nov 2002 Reflectivity

Radar



WSM6

Dataset: ripysu RIP: rip myplots Init: 1200 UTC Sun 10 Nov 02
Fest: 12.00 Valid: 0000 UTC Mon 11 Nov 02 (1800 CST Sun 10 Nov 02)
Max Reflectivity

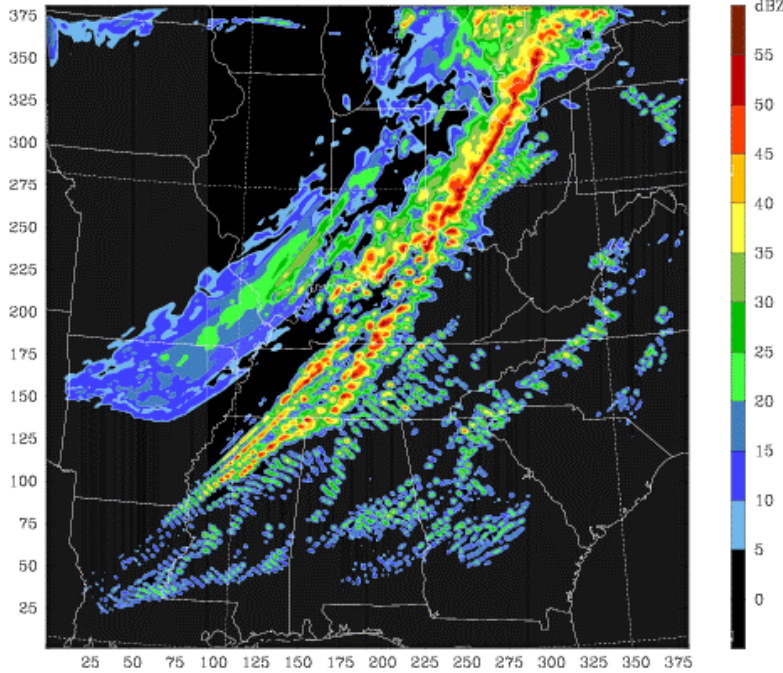


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WSM6

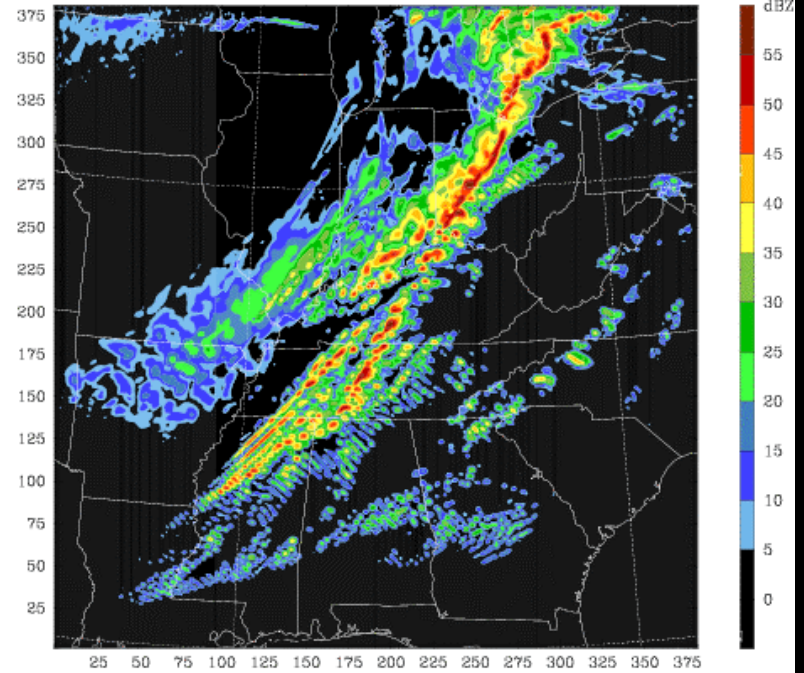
LFO

Dataset: ripysu RIP: rip myplots Init: 1200 UTC Sun 10 Nov 02
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Max Reflectivity



Model info: V2.0.2 No Cumulus YSU PHL WSM 6class 4.0 km, 34 levels, 24 sec

Dataset: ripysulin RIP: rip myplots Init: 1200 UTC Sun 10 Nov 02
Fest: 12.00 Valid: 0000 UTC Mon 11 Nov 02 (1800 CST Sun 10 Nov 02)
Max Reflectivity



Model info: V2.0.2 No Cumulus YSU PHL Lin et al 4.0 km, 34 levels, 24 sec

Noah Land Surface Model

Developers: Fei Chen (NCAR/RAP), Ken Mitchell (NCEP), Mike Ek (NCEP), Mukul Tewari (RAP), and others

- New unified version of Oregon State University (MM5) scheme and NCEP's Eta/LDAS scheme
- Snow-cover fraction
- Frozen soil physics
- Other changes, including emissivity and urban effects

RUC Land Surface Model

Developers: Tanya Smirnova (NOAA/FSL)

- Operational version from RUC
- 6 sub-soil layers
- Multi-layer snow model

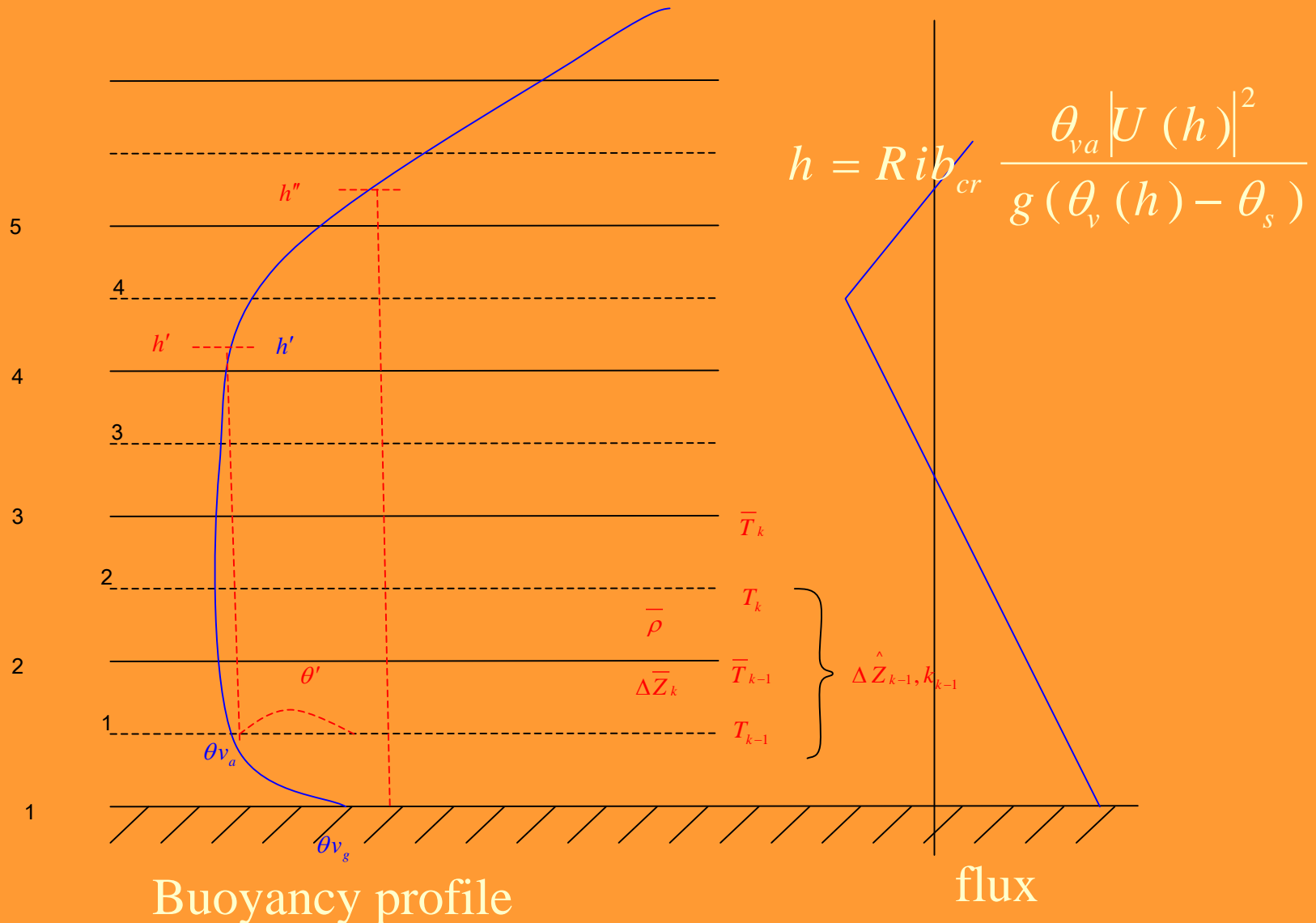
Yonsei University (YSU)

Planetary Boundary Layer

Developers: Song-You Hong and Yign Noh
(YSU)

- Successor to MRF PBL (Hong and Pan)
- Explicit treatment of entrainment layer
- Based on Large-Eddy Model results
- PBL height is lower because it excludes upper part of entrainment layer

MRFPBL (Troen and Mahrt) represents the entrainment **implicitly**
 YSUPBL (Hong and Noh) represents the entrainment **explicitly**



TM Model

New Model

- heat flux profile

i) $z < h$

$$-\overline{w'\theta'} = K_h \left(\frac{\partial\theta}{\partial z} - \gamma_h \right)$$

$$\gamma_h = b_0 \frac{\overline{w'\theta'_0}}{w_{s0}h}$$

ii) $z > h$

not defined

- h is above the height of the minimum

heat flux here

i) $z < h$

$$-\overline{w'\theta'} = K_h \left(\frac{\partial\theta}{\partial z} - \gamma_h \right) - \overline{w'\theta'_h} \left(\frac{z}{h} \right)^3$$

$$\gamma_h = b \frac{\overline{w'\theta'_0}}{w_s(h/2)h}$$

ii) $z > h$

$$-\overline{w'\theta'} = K_h \frac{\partial\theta}{\partial z}$$

$$K_h = \frac{-\overline{w'\theta'_h}}{(\partial\theta/\partial z)_h} \exp \left[-\frac{(z-h)^2}{\delta^2} \right]$$

$$(\delta = 0.02h + 0.05w_m^2 / \Delta b)$$

$$* \overline{w'\theta'_h} = -Aw_m^3 / h, \quad (w_m^3 = w_*^3 + 5u_*^3)$$

- h is the height of minimum heat flux

Troen and Mahrt (1986, BLM)

Noh et al. (2002, BLM)

Cold front (10-11 Nov 2002)

- 4 km grid (cloud-resolving)
- YSU PBL compared to MRF PBL
- Showing how different pre-frontal soundings affect frontal convection

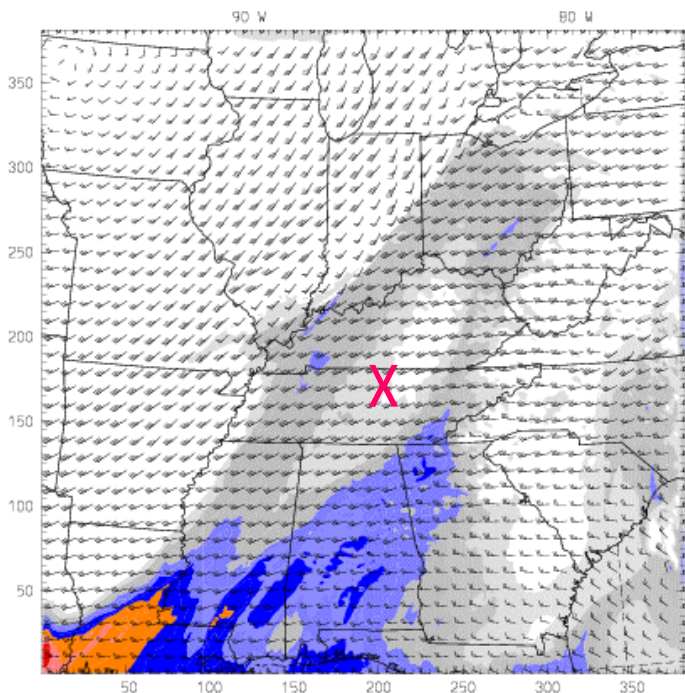
21Z 10 Nov 2002 CAPE

MRF

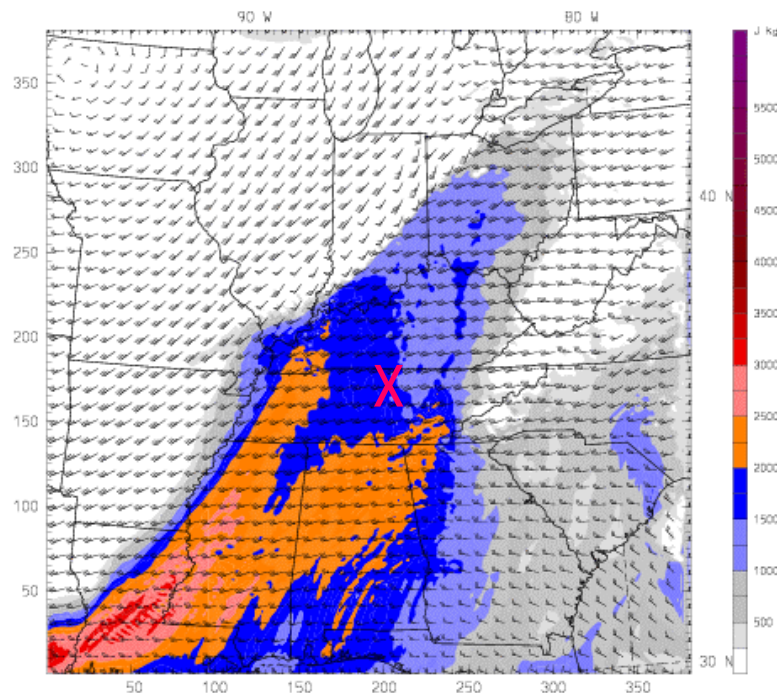
YSU

4km WRF MRF NCAH/MMM
Fcst: 9 h
CAPE (for parcel with max theta-e)
0-6 km shear vectors
Init: 12 UTC Sun 10 Nov 02
Valid: 21 UTC Sun 10 Nov 02 (14 MST Sun 10 Nov 02)

4km WRF YSU NCAH/MMM
Fcst: 9 h
CAPE (for parcel with max theta-e)
0-6 km shear vectors
Init: 12 UTC Sun 10 Nov 02
Valid: 21 UTC Sun 10 Nov 02 (14 MST Sun 10 Nov 02)



BARB VECTORS: FULL BARB = 10 kts
Model info: V1.3.0 No Cumulus MRF PBL Lin 4 km, 34 levels, 25 sec

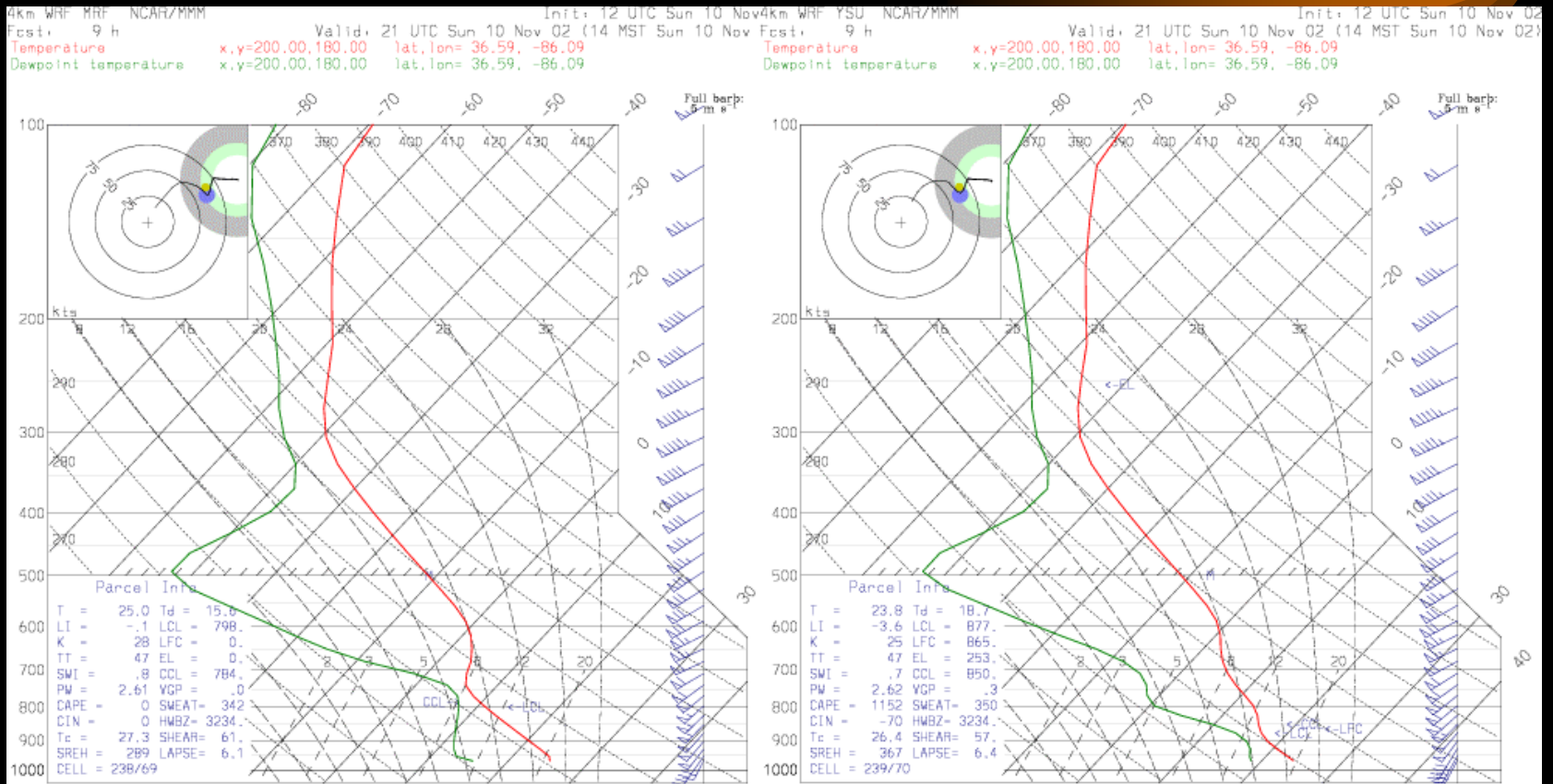


BARB VECTORS: FULL BARB = 10 kts
Model info: V1.3.0 No Cumulus YSU PBL Lin 4 km, 34 levels, 25 sec

21Z 10 Nov 2002 Sounding

MRF

YSU



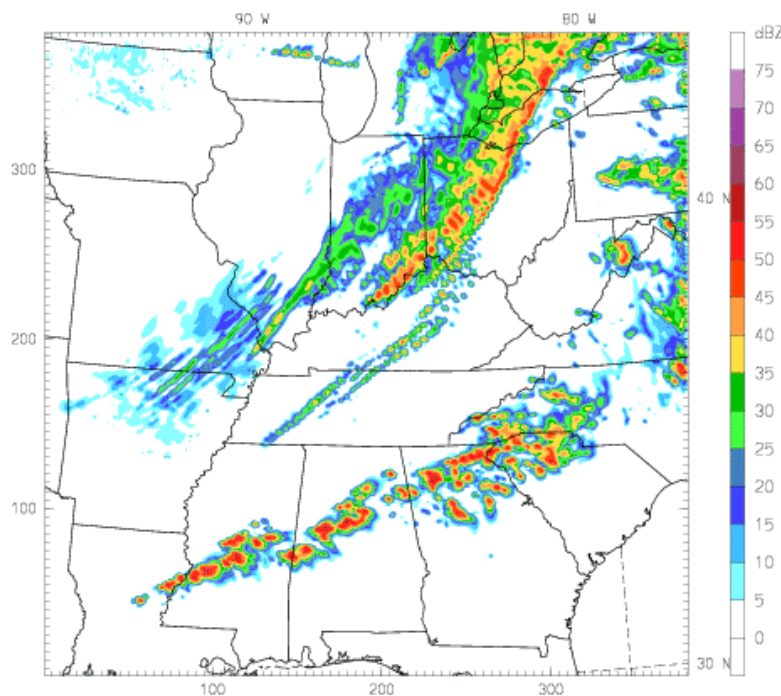
00Z 11 Nov 2002 Reflectivity

MRF

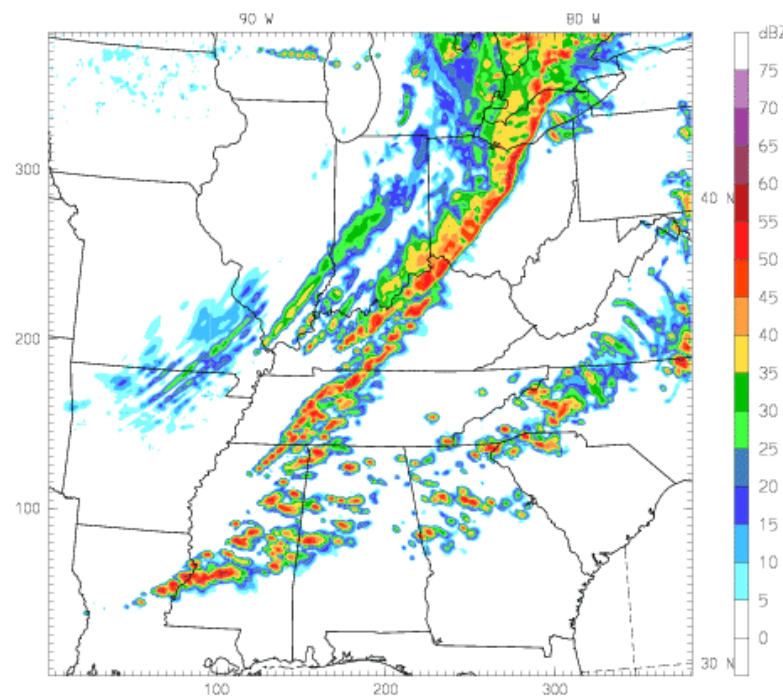
YSU

4km WRF MRF NCAR/MMM
Fcst: 12 h
Max Reflectivity
Init: 12 UTC Sun 10 Nov 02
Valid: 00 UTC Mon 11 Nov 02 (17 MST Sun 10 Nov 02)

4km WRF YSU NCAR/MMM
Fcst: 12 h
Max Reflectivity
Init: 12 UTC Sun 10 Nov 02
Valid: 00 UTC Mon 11 Nov 02 (17 MST Sun 10 Nov 02)



Model info: V1.3.0 No Cumulus MRF PBL Lin 4 km, 34 levels, 25 sec

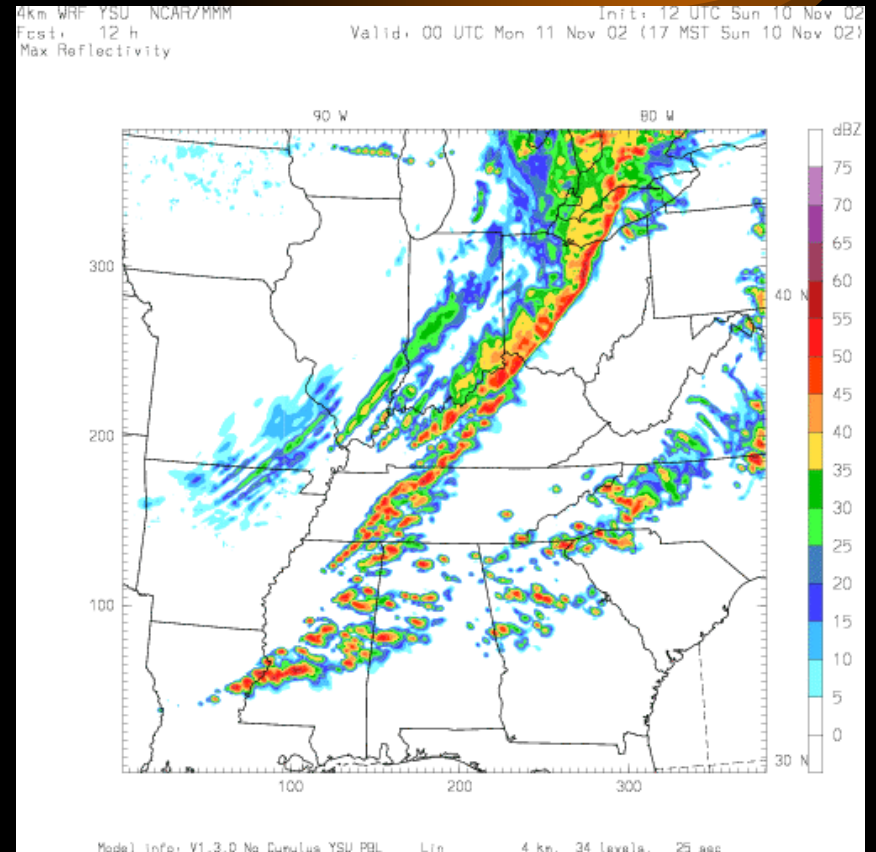
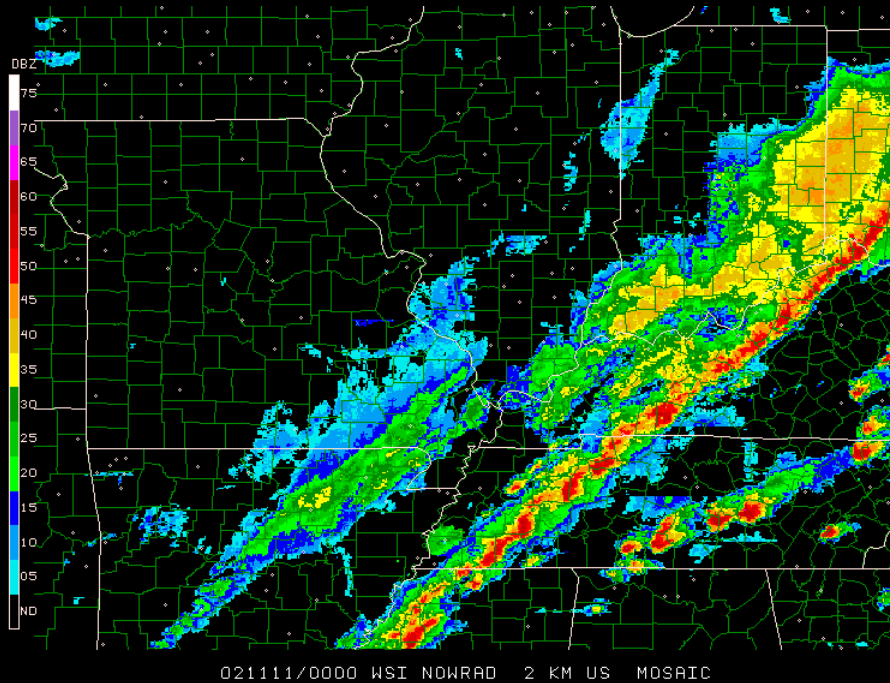


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Radar

YSU



Summary

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